## Images & turbulence – 26

#### -> Detection noises:

• At first: *photon noise* (or *shot noise*), poissonian, actually a transformation of the image.

$$p(n) = \frac{N^n e^{-N}}{n!}$$
, with :  $N = \frac{L\Delta t}{h\nu}$ ,  $L =$ luminosity,  $\Delta t =$ time exp.

p(n) = probability to detect n photons when N are expected

For large N: ~gaussian...

$$p(n) \simeq \exp\left(-\frac{(n-N)^2}{2N}\right)$$

## Images & turbulence – 27

#### -> Detector noises:

• At first: *photon noise* (or *shot noise*), poissonian, actually a transformation of the image.

• At last: *read-out noise* (*RON*), gaussian with zero mean and rms  $\sigma_e$  [e-/px], additive noise.

• In between: *dark current noise, amplification noise* & *exotic dark current noise* in the case of EMCCDs, noise due to the *calibration* of the *flat field, 'salt & pepper' noise* ('hot' and 'cold' pixels), etc.

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; Photon noise (Poisson) if keyword\_set(PHOT\_NOISE) then begin idx=where((image GT 0.) AND (image LT 1E8),c)

- ; For values higher than 1E8, should one if (c NE 0) then for i=01,c-11 do \$ ; really has to worry about photon noise noisy\_image[idx[i]]=randomn(seed\_pn,POISSON=image[idx[i]],/DOUBLE) if
- ; Additive dark-current noise (Poisson)
- if keyword\_set(SIGMA\_DARK) then begin
  - if not(keyword\_set(DELTA\_T)) then begin

message, "dark-current noise calculation does need a time exposure value!!"
endif else noisy\_image+=randomn(seed\_dark,npx,npy,POISSON=sigma\_dark\*delta\_t,/DOUBLE)
endif

;; EMCCD noises

; Additive exotic (time-exposure-independent) dark-current noise (Poisson) if keyword\_set(EXODARK) then noisy\_image+=randomn(seed\_xd,npx,npy,POISSON=exodark,/DOUBLE)

; Additive main EMCCD noise (Gamma)
if keyword\_set(GAIN\_L3CCD) then begin
 idx=where(image GT 0, c)
 if (c NE 0) then for i=0l,c-1l do \$
 noisy\_image[idx[i]]+=gain\_l3ccd\*randomn(seed\_l3ccd,GAMMA=image[idx[i]],/DOUBLE)
; noisy\_image=long(temporary(noisy\_image))
endif

```
;; Flat-field calibration residuals
if keyword_set(FF0FFSET) then begin
   ffres=randomn(seed_ff,npx,npy)*ffoffset+1.
    idx = where(ffres LE 0., c)
    if (c NE 0) then ffres[idx]=1. ; Put possible<=0 ff values to 1.
    noisy_image*=ffres
endif</pre>
```

endif

```
;; Additive read-out noise (Gaussian)
if keyword_set(SIGMA_RON) then $
    noisy_image+=randomn(seed_ron,npx,npy,/NORMAL,/DOUBLE)*sigma_ron
```

```
; Force to zero negative values
if keyword_set(POSITIVE) then begin
    idx=where(noisy_image LT 0, c)
    if (c GT 0) then noisy_image[idx]=0.
endif
```

noisy\_image = addnoise(input\_image,

PHOT\_NOISE=phot\_noise

DELTA\_T=delta\_t, EXODARK=exodark, GAIN\_L3CCD=gain\_l3ccd, FFOFFSET=ffoffset, SIGMA\_RON=sigma\_ron, POSITIVE=positive, OUT\_TYPE=out\_type

#### img formation w/noise:

 'add' photon noise on one short-exp. PSF (in function of N...),
 long-exp. PSF (100N photons!),
 'add' photon noise on the long-exp. PSF,
 compare long-exp. & short-exp. noisy images (and 'clean' images).

### Images & turbulence - 27

[IDL> restore, 'PSF\_r0=10cm\_L0=10m\_lambda=500nm.sav [IDL> help % At \$MAIN\$ CUBE\_PSF FLOAT = Array[128, 128, 100] SHORTEXP DOUBLE = Array[128, 128] Compiled Procedures: \$MAIN\$

Compiled Functions:

[IDL> shortexp=cube\_PSF[\*,\*,0] [IDL> total(shortexp) 0.19702147 [IDL> shortexp=shortexp/total(shortexp)\*100. [IDL> total(shortexp) 99.999664 [IDL> .r addnoise % Compiled module: ADDNOISE. [IDL> shortnoisy=addnoise(shortexp, /PHOT\_NOISE)



comparison: visually or by means of a least mean square distance...

img formation w/noise:

1- 'add' photon noise on one short-exp. PSF (with N photons/img), 2- long-exp. PSF (=> with 100N photons), 3- 'add' photon noise on the long-exp. PSF, 4- compare long-exp. & short-exp. noisy images (and 'clean' images). 5- possibly compare also with stacked image (from the 100 short-exp. noisy images)

### REPORT

```
- Preliminary measures
+ introduction/context
+ PSD(r0, L0)
+ => influence of r0 and L0
+ rms(r0, L0)
+ => influence of r0 and L0
+ FWHM(r0 or lambda=>r0, L0)
+ => influence of r0 and L0
+ => comparison with the "seeing" lambda/r0
+ noisy images
```

# Adaptive optics - 01



## Adaptive optics -02



# Adaptive optics - 03

### Some orders of magnitude concerning AO systems:

	@500nm	@2 <b>.</b> 2µm
spatial sampling (WFS analysis elements size) → d ≈ r <sub>0</sub>	≈ 10 cm	≈ 60 cm
number of WFS analysis elements (≈ number of D → N ∝ (D/d)², with D=10m	M actuators) ≈ 7500	<b>≈ 200</b>
temporal sampling $\rightarrow f \propto 10 v/r_0$	≈ 1 kHz	≈ 0.2 kHz